

Yukawa Unification in $SO(10)$ with light sparticle spectrum

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27th August 2011

based on:

MB, M. Olechowski, S. Pokorski, arXiv:1107.2764 (to appear in JHEP)

- 1 Conditions for $t - b - \tau$ Yukawa unification
- 2 Yukawa unification with $\mu < 0$
 - interplay between BR($b \rightarrow s\gamma$) and $(g - 2)_\mu$
 - implications for SUSY spectrum

Yukawa Unification in SO(10)

Minimal SO(10) models predict unification of the top, bottom and tau Yukawa couplings at M_{GUT}

$b - \tau$ Yukawa unification leads to tree level bottom mass bigger than the observed mass

SUSY loop correction to the bottom mass must be negative:

$$\left(\frac{\delta m_b}{m_b}\right) \approx \frac{g_3^2}{6\pi^2} \mu m_{\tilde{g}} \tan \beta I(m_{\tilde{b}_1}^2, m_{\tilde{b}_2}^2, m_{\tilde{g}}^2) + \frac{h_t^2}{16\pi^2} \mu A_t \tan \beta I(m_{\tilde{t}_1}^2, m_{\tilde{t}_2}^2, \mu^2)$$

For $\mu > 0$

- The gluino-sbottom contribution has wrong sign so has to be suppressed \Rightarrow multi-TeV squarks
- The chargino-stop contribution must dominate $\Rightarrow A_t$ negative and very large to account for the observed bottom mass

Baer et al. '2008

For $\mu < 0$

- The gluino-sbottom contribution has correct sign
 \Rightarrow squark masses may be small
 \Rightarrow The sign and the magnitude of A_t much less restricted

SM prediction for $a_\mu \equiv (g-2)_\mu/2$ is more than 3σ below experimental result:

$$\delta a_\mu = a_\mu^{\text{exp}} - a_\mu^{\text{SM}} = (28.7 \pm 8) \cdot 10^{-10}$$

Dominant SUSY contribution $\sim \text{sgn}(\mu M_2) \tan \beta$

For phenomenologically acceptable models:

$$\mu < 0 \quad \Rightarrow \quad M_2 < 0$$

Is it natural to consider $M_2 < 0$ in the context of $SO(10)$ GUT?

YES \Rightarrow non-singlet F -terms give non-universal gaugino masses at M_{GUT} .

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Gaugino masses in SUGRA can arise from dimension 5 operator:

$$\mathcal{L} \supset -\frac{F^{ab}}{2M_{\text{Planck}}}\lambda^a\lambda^b + \text{c.c.}$$

- $\langle F^{ab} \rangle$ must transform as a singlet under the SM gauge group
- $\langle F^{ab} \rangle$ must belong to the symmetric part of $\mathbf{45} \times \mathbf{45}$
- $\langle F^{ab} \rangle$ can be in a non-singlet representation of $\text{SO}(10)$

$$(\mathbf{45} \times \mathbf{45})_S = \mathbf{1} + \mathbf{54} + \mathbf{210} + \mathbf{770}$$

If $\langle F^{ab} \rangle$ transforms as $\mathbf{54}$, gaugino masses are given by:

$$M_1 : M_2 : M_3 = -\frac{1}{2} : -\frac{3}{2} : 1$$

Martin, 2009

$M_2 < 0 \Rightarrow \mu < 0$ can be consistent with $(g-2)_\mu$ constraint!

At large $\tan\beta$ proper REWSB requires $m_{H_d}^2 - m_{H_u}^2 \gtrsim M_Z^2$

For universal soft terms at the GUT scale:

- RG evolution results in positive (negative) contribution to $m_{H_d}^2 - m_{H_u}^2$ proportional to $M_{1/2}^2$ (m_0^2)
- REWSB possible only if $M_{1/2} > m_0$
- μ^2 strongly correlated with $M_{1/2}^2$
 \Rightarrow too large SUSY threshold correction to the bottom mass

Carena, Olechowski, Pokorski, Wagner '1994

proper REWSB requires e.g. Higgs splitting at M_{GUT} : $m_{H_u} < m_{H_d}$

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Non-universal scalar masses in SO(10)

All sfermions in **16**-dim rep. of SO(10) while Higgses in **10**-dim rep.
⇒ Pattern of soft scalar masses restricted by SO(10) gauge symmetry:

$$m_{H_d, H_u} = m_{10}$$
$$m_{Q, U, D, L, E} = m_{16}$$

This is not enough for $t - b - \tau$ Yukawa unification with proper REWSB

D-term contribution

Rank of SO(10) is larger than the rank of SM gauge group
⇒ in the effective theory below the GUT scale soft scalar masses acquire new contribution proportional to D -term and charges of the broken U(1):

Kawamura, Murayama, Yamaguchi '1994

$$m_{H_d}^2 = m_{10}^2 + 2D$$
$$m_{H_u}^2 = m_{10}^2 - 2D$$
$$m_{Q, U, E}^2 = m_{16}^2 + D$$
$$m_{D, L}^2 = m_{16}^2 - 3D$$

$D > 0$ may allow for proper REWSB

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$$\mu < 0$$

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$$m_{D,L}^2 = m_{16}^2 - 3D$$

- Non-universal gaugino masses:

$$M_1 = -\frac{1}{2}M_{1/2}$$

$$M_2 = -\frac{3}{2}M_{1/2}$$

$$M_3 = M_{1/2}$$

- Universal trilinear couplings: $A_U = A_D = A_E = A_0$

5 parameters + $\tan \beta$

$(g - 2)_\mu$ vs $b \rightarrow s\gamma$

$(g - 2)_\mu$

SM prediction is more than 3σ below the experimental result
 \Rightarrow **light** SUSY spectrum is required for significant (positive) SUSY contribution

$\text{BR}(b \rightarrow s\gamma)$

SM prediction is in 1σ agreement with the experimental data
Charged Higgs contribution increases SM result
 \Rightarrow **heavy** SUSY spectrum is preferred

How to disentangle $(g - 2)_\mu$ from $b \rightarrow s\gamma$?

- 1 $\tilde{\chi}^\pm - \tilde{\nu}_\mu$ contribution:

$$a_\mu^{\chi^\pm} \approx \frac{g_2^2}{32\pi^2} \frac{m_\mu^2}{M_{\text{SUSY}}^2} \text{sgn}(\mu M_2) \tan \beta$$

- 2 $\tilde{\chi}^0 - \tilde{\mu}$ contribution:

$$a_\mu^{\chi^0} \approx \frac{1}{192\pi^2} \frac{m_\mu^2}{M_{\text{SUSY}}^2} [g_1^2 \text{sgn}(\mu M_1) - g_2^2 \text{sgn}(\mu M_2)] \tan \beta$$

Our model: $m_{\tilde{\nu}_\mu} \approx m_{\tilde{\mu}_L} < m_{\tilde{\mu}_R}$ (due to D -term splitting)



chargino contribution dominates

SUSY contributions to $b \rightarrow s\gamma$

Chargino - squarks contribution enhanced at large $\tan\beta$:

$$C_{7,8}^X \supset s_{\tilde{t}} c_{\tilde{t}} \frac{\tilde{U}_{a2} \tilde{V}_{a2} m_t}{\sin 2\beta m_{\chi_a^+}} \left[F_{7,8}^{(3)} \left(\frac{m_{\tilde{t}_1}^2}{m_{\chi_a^+}^2} \right) - F_{7,8}^{(3)} \left(\frac{m_{\tilde{t}_2}^2}{m_{\chi_a^+}^2} \right) \right] \\ - \frac{\tilde{U}_{a2} \tilde{V}_{a1} M_W}{\sqrt{2} \cos \beta m_{\chi_a^+}} \left[F_{7,8}^{(3)} \left(\frac{m_{\tilde{q}}^2}{m_{\chi_a^+}^2} \right) - c_{\tilde{t}}^2 F_{7,8}^{(3)} \left(\frac{m_{\tilde{t}_1}^2}{m_{\chi_a^+}^2} \right) - s_{\tilde{t}}^2 F_{7,8}^{(3)} \left(\frac{m_{\tilde{t}_2}^2}{m_{\chi_a^+}^2} \right) \right]$$

Degrassi, Gambino, Giudice '2000

Signs (relative to SM contribution) are the same as:

$\text{sgn}(\mu A_t)$ for the **stop-mixing part**; $\text{sgn}(-\mu M_2)$ for the **"gaugino" part**

Chargino contribution should be negative to satisfy $b \rightarrow s\gamma$

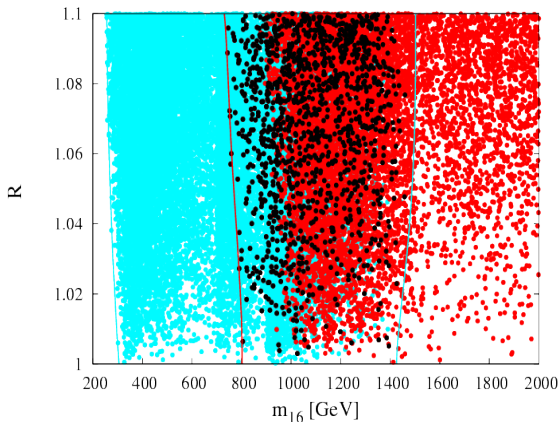
A_t is often driven to negative values by RGEs and the **stop-mixing part** has wrong (positive) sign for $\mu < 0$

"gaugino" part is always negative (as preferred by phenomenology)

BR($b \rightarrow s\gamma$) prefers:

- the lightest chargino is gaugino-like ($\mu^2 > M_2^2$)
- small or positive A_t
- large mass splitting between stops and 1st/2nd generation of squarks

$$R \equiv \frac{\max(h_t, h_b, h_\tau)}{\min(h_t, h_b, h_\tau)} \Big|_{\text{GUT}}$$



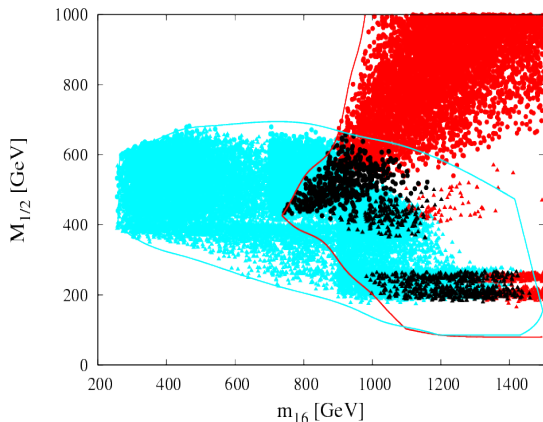
$(g - 2)_\mu$ satisfied at 2σ

$b \rightarrow s\gamma$ satisfied at 2σ

$b \rightarrow s\gamma$ and $(g - 2)_\mu$
satisfied at 2σ

$t - b - \tau$ Yukawa unification consistent with $b \rightarrow s\gamma$ and $(g - 2)_\mu$

The WMAP bound on Ω_{LSP} is also important



$$R \leq 1.1$$

$(g - 2)_\mu$ satisfied at 2σ

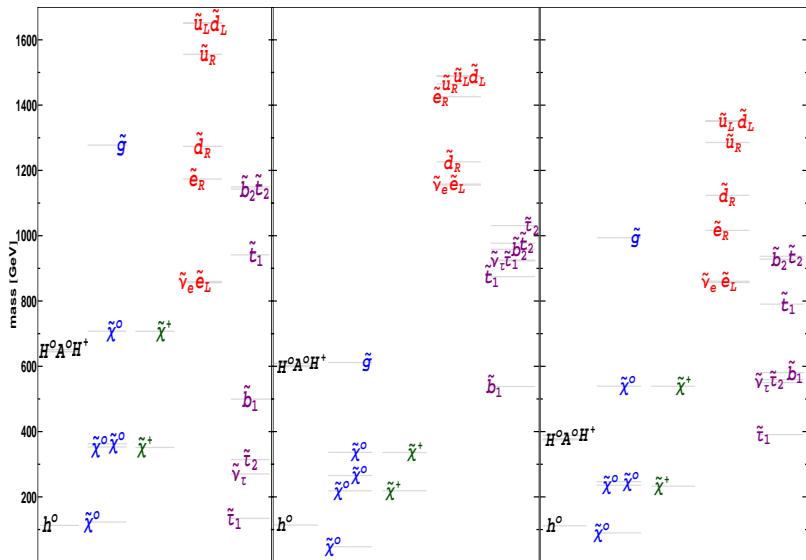
$b \rightarrow s\gamma$ satisfied at 2σ

$b \rightarrow s\gamma$ and $(g - 2)_\mu$
satisfied at 2σ

The WMAP bound selects two allowed regions of $M_{1/2}$ (and gluino mass):

- resonant annihilation through Z or $h^0 \Rightarrow$ gluino mass $\sim 500 - 700$ GeV
- stau co-annihilations, A -funnel \Rightarrow gluino mass $\sim 900 - 1600$ GeV

Typical spectra



Conclusions

- $t - b - \tau$ Yukawa coupling unification much more natural for $\mu < 0$
- Non-universality of soft terms at M_{GUT} are necessary for proper REWSB
- Appropriate pattern of non-universalities easily accommodated in SUSY SO(10) GUT:
 - $m_{10} - m_{16}$ splitting and D -term contribution to scalar masses
 - gaugino masses from non-singlet F -term in **54** of SO(10)
- $(g - 2)_\mu$ anomaly explained without violating $b \rightarrow s\gamma$
- bino-like LSP (with small admixture of higgsino) may play a role of dark matter in the Universe:
 - stau coannihilations,
 - resonant annihilations through h^0 , Z or A^0
- Light SUSY spectrum:
 - the whole SUSY spectrum easily below 1.5 TeV
 - predicted gluino mass is 500-700 GeV or 900-1600 GeV